

Chapter 1

Basic Properties of Light

Light is described using 3 pictures - seemingly contradictory!

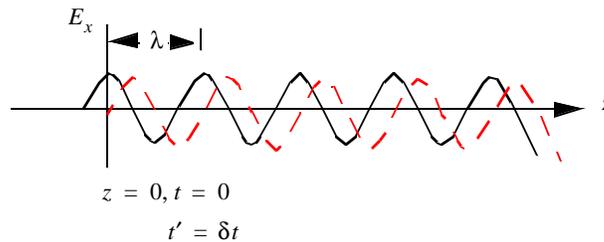
Waves - Rays - Photons

Waves

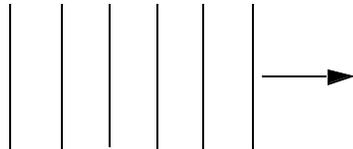
[Reading Assignment: Hecht, Chapter 2 (most of this should be review), 3.2, 3.3, 3.4.4, 3.5, 3.6]

A propagating “disturbance” in *electric* and *magnetic* field (simultaneously!)

Example:
$$\vec{E} = E\hat{x}\cos\left(\frac{2\pi}{\lambda}z - 2\pi ft\right)$$



- At a fixed point in space, the electric field oscillates in time. At a fixed point in time, we see a wave train frozen.
- This is called a plane-wave because the field is constant everywhere in the x-y plane at a given z. Another way to draw this is



“wave-fronts” surface of constant “phase” or “phase-fronts”

- The wavefront advances by a distance λ , in a time $1/f$. So the velocity is $v = \text{distance/time} = \lambda f$. One of the many remarkable properties of light is it’s universal, constant speed:

$$c = 2.997 \times 10^8 \text{ m/sec}$$

$$c = \lambda f \text{ in vacuum}$$

- The physics of electromagnetic (EM) wave propagation is valid for arbitrary λ, f . On Earth, we can generate, manipulate and/or detect EM waves with wavelength from ~ 100 km all the way down to $\sim 10^{-6}$ A°. Usually we describe light by wavelength rather than frequency, except in the microwave and radio regions.

- The electromagnetic spectrum encompasses the complete range of frequency/wavelength. Different regions have different names. Radio, microwave, infrared, visible, ultraviolet, x-ray, γ -ray.

Index of Refraction

- When light travels in materials, the speed is modified:

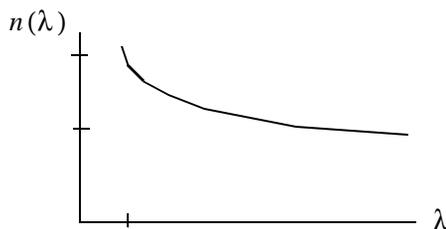
$$v = \frac{c}{n} = \lambda f$$

Usually $n \geq 1$. (It can be < 1)

- The reason is that the electric field shakes the electrons, which tends to drag the field.
- Plane wave still has the same form: $\vec{E} = E\hat{x}\cos\left(\frac{2\pi}{\lambda}z - 2\pi ft\right)$, but the effective wavelength becomes modified by n . If we define the vacuum wavelength. $\lambda_{\text{vac}} = \frac{c}{f}$ then in the material $\lambda = \frac{c}{nf} = \frac{\lambda_{\text{vac}}}{n}$. The wavelength becomes shorter, if $n > 1$.

Dispersion

The index of refraction in most materials depends on wavelength. $n(\lambda)$. This is called dispersion.



In air – the index depends also on air pressure, humidity, and temperature which leads to many beautiful atmospheric effects.

Wavelength units (length)

We commonly use Angstrom units (\AA) for light wavelength.

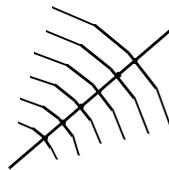
$$1 \text{\AA} = 10^{-8} \text{ cm} = 0.1 \text{ nm}$$

This is of the order of the size of an atom. We also use standard metric units: m, cm, mm, nm

Visible light $\sim 4000 \rightarrow 7000 \text{ \AA}$, $400 \rightarrow 700 \text{ nm}$, $0.4 \rightarrow 0.7 \text{ }\mu\text{m}$

Spherical Waves

Another type of ideal light wave. Constant phase fronts are circular, emanating from a point source. Far away from the source, the radius of the circle becomes so large that we can approximate the wave as a plane wave.



For spherical waves, we have

$$E = E_0 \frac{\cos\left(\frac{2\pi}{\lambda} r - 2\pi f t\right)}{r}$$

Huygens' Principle

Very useful model for wave propagation.

- Every point on a wavefront is regarded as a secondary point source generating a spherical wavelet.
- The advance of the wave front is found at the envelope of all these wavelets



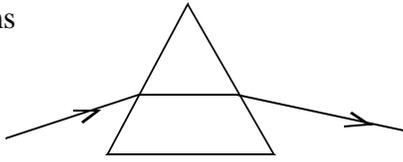
- Generally, this seems to give parallel wavefronts. But things get interesting at edges. This leads to diffraction (more later).

Rays

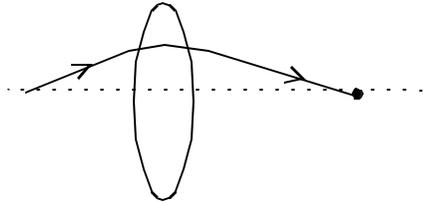
- Follow a point on the wavefront. As the wavefront advances the point traces a straight line. This is a ray of light.

- For many cases, we can forget the waves and just trace rays in optical systems. This allows a vast simplification of our analysis and design processes. Virtually all optical design is done with rays. Highly sophisticated optical design CAD programs are available for ray tracing.

Prisms



Lenses



Photons (light “particles”)

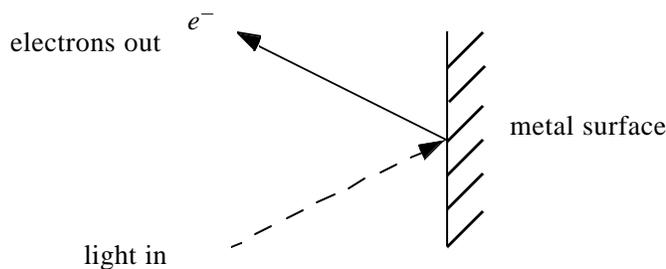
- This picture has light represented by tiny bundles of energy (or quanta), following straight line paths along the rays.
- The coexistence of electromagnetic wave physics and photon physics is the central paradox of quantum mechanics.
- Each photon has an energy given by

$E = h\nu$ $h = 6.62 \times 10^{-34} \text{ J}\cdot\text{s}$ $1 \text{ W} = 1 \text{ J per second}$	h is Planck's constant
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For 2 eV visible photons, <table border="1" style="border-collapse: collapse; display: inline-table; margin-left: 20px;"> <tr> <td style="padding: 5px;"> $1 \text{ W} = 6.3 \times 10^{18} \text{ eV/s}$ $= 3.15 \times 10^{18} \text{ photons/sec}$ </td> </tr> </table>	$1 \text{ W} = 6.3 \times 10^{18} \text{ eV/s}$ $= 3.15 \times 10^{18} \text{ photons/sec}$
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- Light power \rightarrow photons/sec

Photoelectric Effect

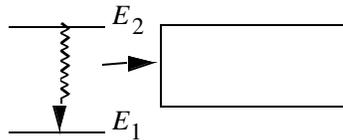


- The electron energy is directly related to the photon energy.

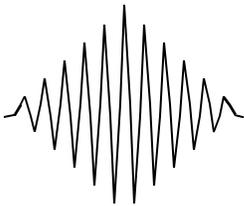
- When the photon energy is below threshold value, no electrons are emitted . The threshold depends on the metal. It is called the work function.
- When the light power is low $\sim 10^2 \rightarrow 10^3$ photons/sec \rightarrow each individual electron can be separated and counted. This called photon counting (more later).

Atomic Radiation

- Atoms have energy states corresponding to electron orbits.



- One atom “jumps” from a higher energy state to a lower energy state and emits one photon.
- Photons are not point particles. They have a wave-like property. A useful picture is the wave-packet.



Many photon packets can be thought of as superimposing to make up a plane wave, spherical wave or any other wave.

The typical photon energy unit is the electron-Volt. This is defined as the energy required to push one electron across a one-Volt potential,

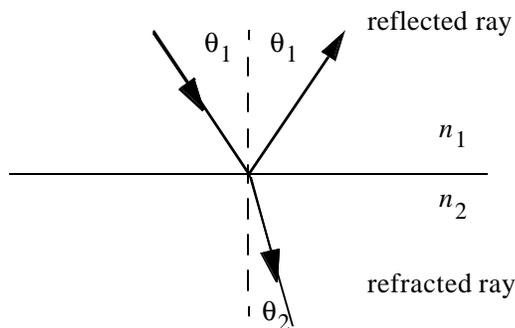
$$1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}$$

Typical visible photon energy $\sim 1.2 \rightarrow 2.3 \text{ eV}$

Reflection and Refraction, Snell's Law

[Reading assignment: Hecht 4.3, 4.4, 4.7]

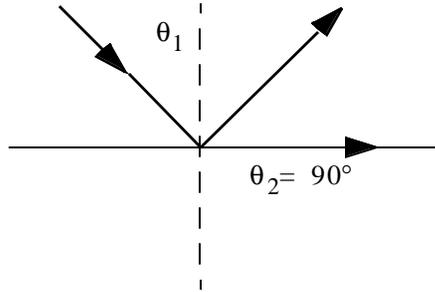
An important element of optics is the interface between 2 materials with different index of refraction



$$\frac{\sin\theta_1}{\sin\theta_2} = \frac{n_2}{n_1} \rightarrow \begin{array}{ll} \text{if } n_2 > n_1 & \theta_2 < \theta_1 \\ \text{if } n_1 > n_2 & \theta_2 > \theta_1 \end{array}$$

Total Reflection

If $n_1 > n_2$, then we can have

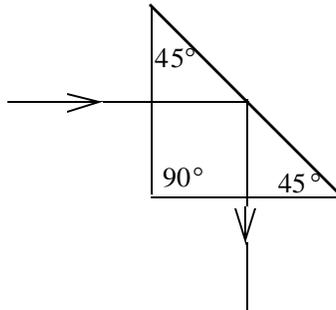


$$\begin{aligned} \theta_2 &= 90^\circ \\ \sin\theta_c &= \frac{n_2}{n_1} \\ \theta_c &= \sin^{-1}(n_2/n_1) \end{aligned}$$

The refracted ray disappears! The light is totally reflected. This usually occurs inside a prism, and is called total internal reflection. θ_c = “critical angle”. For a typical glass with $n= 1.5$, the critical angle is:

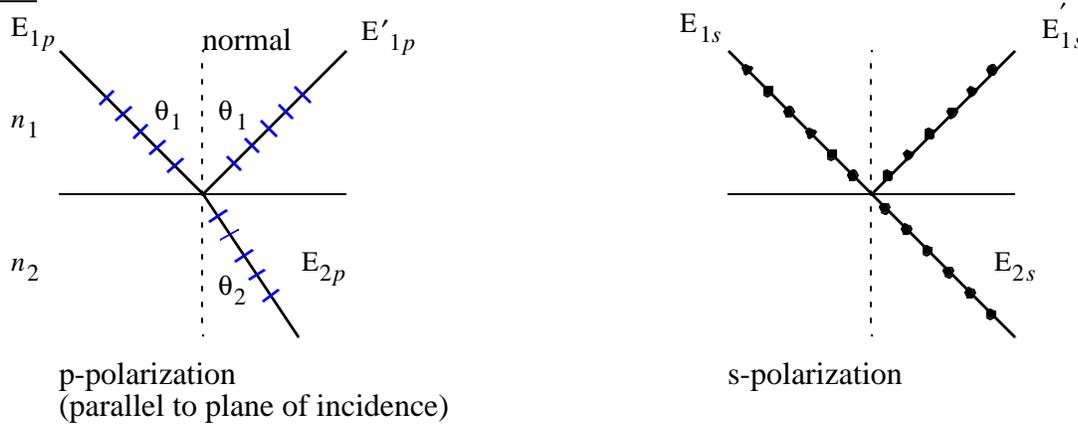


So for $\theta= 45^\circ$, the light is reflected. A very common prism is the right angle prism



Light Impinging at a Surface

The plane containing the light ray propagation vector and the surface normal is called the “plane of incidence”



- For a general polarization state incident on a surface, we choose s and p directions to decompose the polarization effects.

Fresnel Reflection Coefficients

[Optional reading: Hecht 4.6]

The magnitude of reflection and transmission at an interface between n_1 and n_2 are given by

$$R_p = \frac{\tan^2(\theta_2 - \theta_1)}{\tan^2(\theta_2 + \theta_1)} \quad T_p = \frac{n_2 \cos \theta_2}{n_1 \cos \theta_1} \frac{4 \sin^2 \theta_2 \cos^2 \theta_1}{\sin^2(\theta_1 + \theta_2) \cos^2(\theta_1 - \theta_2)}$$

$$R_s = \frac{\sin^2(\theta_2 - \theta_1)}{\sin^2(\theta_2 + \theta_1)} \quad T_s = \frac{n_2 \cos \theta_2}{n_1 \cos \theta_1} \frac{4 \sin^2 \theta_2 \cos^2 \theta_1}{\sin^2(\theta_1 + \theta_2)}$$

near $\theta_1 = 0$, $R_s = R_p = \left(\frac{n_2 - n_1}{n_2 + n_1} \right)^2$. Notice also, if

$$\theta_2 + \theta_1 = \frac{\pi}{2}, \text{ then } \tan \rightarrow \infty, \text{ and } R_p \rightarrow 0.$$

θ_1 can be shown to be given by . This angle is known as “Brewster’s angle” or the “Polarizing angle”.

Polarization

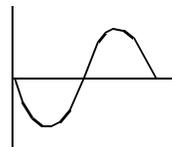
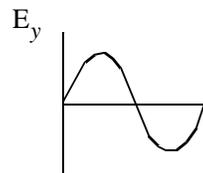
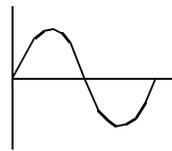
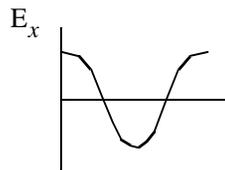
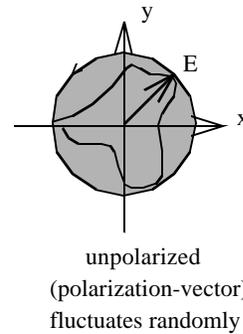
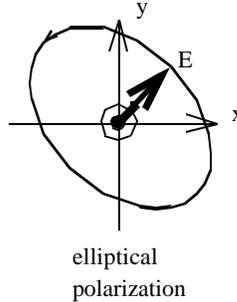
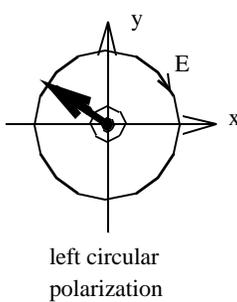
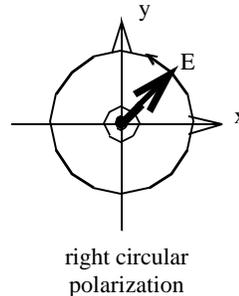
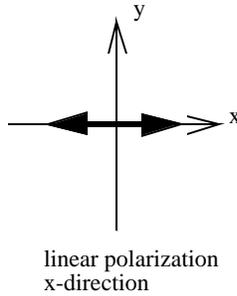
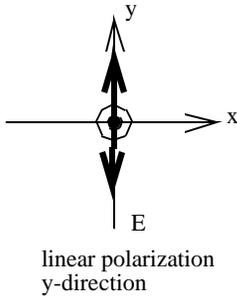
Reading Assignment: Hecht 8.1, 8.2]

- Light waves have transverse polarization

- The electric field vector points in a direction perpendicular to the propagation direction (ray direction).
- The magnetic field vector is orthogonal to propagation direction. Generally, we can ignore the magnetic field.
- The e-field vector can lie anywhere in transverse plane

Polarization State

- The e-field oscillates in time at a given point in space
- For light wave propagating in z-direction, let's look in the x-y plane.



Circular

Linear, 45°

Polarizers are devices which select one polarization

- Polarizing sheet has an allowed direction
 - transmits polarization component of incident light along the allowed direction
 - transmitted light is linearly polarized
- polarizing beamsplitter

